# Development of GPS Constellation Power Monitor System for High Accuracy Calibration/Validation of CYGNSS L1B Data

Tianlin Wang<sup>1</sup>, Christopher S. Ruf<sup>1</sup>, Scott Gleason<sup>2</sup>, Bruce Block<sup>1</sup>, Darren McKague<sup>1</sup>, Damen Provost<sup>1</sup> <sup>1</sup>Department of Climate and Space Sciences and Engineering, University of Michigan, Ann Arbor, MI <sup>2</sup>Southwest Research Institute. Boulder, CO







### **Cyclone Global Navigation Satellite System**

- Constellation of 8 LEO satellites
- GNSS-R bi-static configuration
- Measure ocean surface wind speed
  - In inner core of tropical cyclone
  - With extremely short revisit time
- Launched on December 15, 2016
- · Operating on Science Mode



Figure 1 Bi-static configuration

Table 1 CYGNSS science data products

Level	Description
0	Unprocessed DDMs and metadata
1A	Decompressed, calibrated DDMs, power in Watts
1B	Calibrated DDMs, bistatic radar cross section
2A	Time tagged wind speed of a 25×25 km cell centered on the specular point
2B	Time tagged mean square slope of a 25×25 km cell centered on the specular point
3A	Wind Speed, gridded in time and space
3B	Wind Speed optimized for observing system experiment data assimilation

#### **Global Positioning System (GPS)**

- Constellation of 32 MEO satellites
- Orbital height 12,540 miles (~12 hours period)

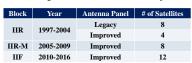




Figure 2 GPS constellation

### **Transmit Power and Antenna Patterns of GPS**

- Transmit power of GPS satellite is about 25 watts.
- IIR/IIR-M built by Lockheed Martin and IIF built by Boeing.
- Only 12 IIR and 8 IIR-M GPS satellites' antenna patterns are published.

Pattern Asymmetry and GPS Yaw Attitude

• The antenna pattern of GPS satellites are azimuthally asymmetric.

Yawing of GPS satellites affects the gain of transmit antenna.

• For  $\theta = \pm 12^{\circ}$ , the difference between can be as large as 0.5 dB.



Figure 5 Satellite vawing



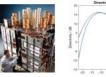


Figure 6 SV coordinate system Figure 7 Pattern diff. map

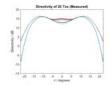


Figure 3 GPS satellite, legacy panel, and improved panel Figure 4 Azimuth averaged patterns

#### **Research Objectives and Science Goal**

- Develop a ground-based GPS constellation power monitor (GCPM) system
- Concurrently monitor the GPS yaw attitude and assess its effect on the received power.
- Retrieve transmit power and antenna pattern of all GPS satellites.
- · Accurately and precisely characterize GPS EIRP for CYGNSS L1B calibration algorithm to improve accuracy and reliability of Level 1 BRCS data
- Provide better estimate of ocean surface wind speed

# Forward Model and Retrieval Algorithm

Forward model is based on Friis radar equation

$$\frac{P_r}{P_t} = G_t(\theta_t, \phi_t) G_r(\theta_r, \phi_r) \left(\frac{\lambda}{4\pi r}\right)^2 L_A$$

- The objective is to retrieve  $P_t$  and  $G_t(\theta_t, \phi_t)$  from the time series received signal.
- Initial value: transmit power = 25 watts; antenna pattern = 5th order polynomial fitting of azimuthally averaged radiation patterns from Lockheed Martin.
- Retrieval algorithm: The minimum-squared-error solution is that value which minimizes the squared difference between  $\bar{y}$  and  $F(\bar{x})$ ,  $\bar{x}_{MSE} = min \{|\bar{y} - F(\bar{x})|^2\}$ .  $\bar{x}_{MSE}$  is estimated iteratively  $\vec{x}_{i+1} = \vec{x}_i + (J^T J + R)^{-1} J^T [\bar{y} - F(\vec{x}_i)]$

### GPS Constellation Power Monitor



- The antenna and thermal box for LNA and calibration system are mounted on the roof top of Space Research Building, University of Michigan, Ann Arbor.
- The Septentrio GPS receiver and GCPM computer are located in GPS Moldwin Lab.
- The processed data will be forwarded to the Science Operation Center of CYGNSS.
- Javad RingAnt-DM antenna is used to receive GPS signals. Full radiation pattern measured at OSU. Elevation mask is set to 20° to filter out the multipath signal.







Figure 9 Javad RingAnt-DM: antenna pattern, location, and panoramic view

#### **Scaled Simulation and Uncalibrated Measurement**

- The measurement power is computed as I<sup>2</sup>+O<sup>2</sup>. The simulation data is computed from the forward model. A scale factor is applied to the simulated data.
- Simulation and measurement matches best with corrected azimuth angle.
- Strong peak due to increase of LNA gain at sunrise (Local time 8AM).

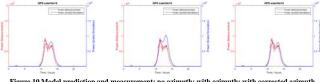


Figure 10 Model prediction and measurement: no azimuth: with azimuth: with corrected azimuth

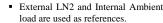
## Thermal Box with LNA&Calibration Subsystem

- · LNA and calibration subsystem are designed and implemented on a PID controlled thermal baseplate.
- · System bandwidth is 5 MHz centered at 1.57542 GHz (GPS L1).
- Rabbit SBC is used to control and record system states.
- . It has solar reflective white coating.



#### **Radiometric Calibration**

- 6 states: EXT LN2 Start, INT Cold. INT Cold+ND, INT Ambient, INT Ambient+ND, EXT LN2 End.
- Measure each state for 0.5 hour. One routine gives 3 hour data. Do 3 routines



- Use Maury noise calibration systems MT7000 series.
- TB of LN2 cold load is 80.5 K. Equivalent TB of LN2 load at K2 is 120.7666 K.

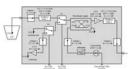




Figure 13 Liquid nitrogen calibration

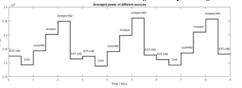


Figure 14 Averaged power of different sources

Table 2 Calibrated brightness temperature of internal sources

Source	Brightness temperature (K)
Internal cold load	59.3937
Excess Noise diode (TB <sub>Ambient+ND</sub> - TB <sub>Ambient</sub> )	110.8693
Excess noise diode (TB <sub>Cold+ND</sub> - TB <sub>Cold</sub> )	110.1614

#### Conclusion

- · Yaw attitude and maneuver of GPS satellites do affect the received power from GPS satellites because of the antenna pattern asymmetry.
- The radiometric calibration system has been designed and implemented. The liquid nitrogen calibration is performed.

#### **Future Work**

- Move the thermal box to the roof and calibrate the received power.
- Include yaw attitude data from GIPSY-OASIS to the forward model.
- Incorporate the retrieved GPS parameters to CYGNSS L1B calibration.
- Error analysis of GCPM and CYGNSS L1B data.

# Acknowledgements

The authors would like to acknowledge the significant contributions and technical support from Space Physics Research Laboratory, University of Michigan and ElectroScience Laboratory, the Ohio State University.